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**Dany Vanbeveren,W. van Rensbergen,C. de
Loore**

NASA Newsletter for International Ultraviolet Explorer (IUE). ,1980

International Ultraviolet Explorer (IUE) NASA Newsletter ,1982

Instability and Variability of Hot-Star Winds Anthony F.J. Moffat, Stanley P. Owocki, Alex W. Fullerton, Nicole St-Louis, 2012-12-06 This specialized workshop was conceived during the workshop on Non isotropic and Variable Outflows from Stars, which was held at the Space Telescope Science Institute in October, 1991. At that meeting, the four of us collectively decided that the time was ripe for an even more focussed discussion of the basic issues in the area of hot-star wind instability and its observable manifestations. Not that the big problems have been solved! Rather, we are currently in a phase of rapid development, both with regard to the models and to the observations. The key issue at this new workshop would be to decide how the time-dependent structures observed in hot-star winds (e. g. , NACs, DACs, blobs, clumps, filaments, shells, puffs, jets, etc.) relate to radiative and other instabilities. Further questions concern the role of turbulence and the nature of its driver, and the effect of stellar rotation, pulsation, and magnetic fields on time-dependent phenomena in hot-star winds. Of no less importance is the impact of stellar wind variability on the derivation of mass-loss rates, on stellar evolution, and on momentum/energy deposition in the interstellar medium. To attain our goal of maximum confrontation (in the positive sense!) we decided: (1) to limit the workshop to the observers and theoreticians most active in this field in the world; (2) to insist that virtually all participants present a talk, thereby avoiding the distraction of poster sessions; and (3) to allocate approximately half of the allotted time

to discussion.

Literature 1987, Part 1 S. Böhme, U. Esser, H. Hefele, I. Heinrich, W. Hofmann, D. Krahn, V. R. Matas, L. D. Schmadel, G. Zech, 2013-11-11 Astronomy and Astrophysics Abstracts aims to present a comprehensive documentation of the literature concerning all aspects of astronomy, astrophysics, and their border fields. It is devoted to the recording, summarizing, and indexing of the relevant publications throughout the world. Astronomy and Astrophysics Abstracts is prepared by a special department of the Astronomisches Rechen-Institut under the auspices of the International Astronomical Union. Volume 43 records literature published in 1987 and received before August 15, 1987. Some older documents which we received late and which are not surveyed in earlier volumes are included too. We acknowledge with thanks contributions of our colleagues all over the world. We also express our gratitude to all organizations, observatories, and publishers which provide us with complimentary copies of their publications. Starting with Volume 33, all the recording, correction, and data processing work was done by means of computers. The recording was done by our technical staff members Ms. Helga Ballmann, Ms. Beate Gobel, Ms. Monika Kohl, Ms. Sylvia Matyssek, Ms. Doris Schmitz-Braunstein, Ms. Uta-Barbara Stegemann. Mr. Jochen Heidt and Mr. Kristopher Polzine supported our task by careful proof reading. It is a pleasure to thank them all for their encouragement. Heidelberg, October 1987

The Editors Contents Introduction 1 Concordance Relation: PHYS-AAA 3 Abbreviations 5 Periodicals, Proceedings, Books, Activities 001 Periodicals 10 002 Bibliographical Publications, Documentation, Catalogues, Data Bases 50 003 Books

Wolf-Rayet Stars Karel A. van der Hucht, Peredur M.

Williams,2012-12-06 In this IAU Symposium on Wolf-Rayet stars, binary aspects received ample attention, notably because of the recognition that many observations of spectral and photometric variability at all accessible wavelengths are related to colliding winds or other forms of wind interaction. The basic structure of the conference and its proceedings is basic parameters and general properties of WR stars; state of the art model atmospheres for WR stars, anisotropic mass loss and disk formation of WR stars, properties of WR binaries; influence of stellar winds on mass transfer in hot massive binary evolution; dust formation near WR stars and other circumstellar phenomena; and hydrodynamics and high-energy physics of colliding winds in WR+O binaries and of WR winds interacting with compact objects. Within this framework 20 invited reviews, 38 invited oral contributions, and 76 poster papers were presented at the Symposium, entertaining 111 astronomers from 24 countries. These proceedings provide up-to-date information on all aspects of Wolf-Rayet atmospheres, binaries, and colliding winds.

American Hereford Record and Hereford Herd Book
American Hereford Cattle Breeders' Association,1895 Brief history of Hereford cattle: v. 1, p. 359-375.

Treatise on the Law of Securities Regulation Thomas Lee Hazen,2009

Board of Contract Appeals Decisions United States. Armed Services Board of Contract Appeals,2001

Wolf-Rayet Stars: Observations, Physics, Evolution C. de Loore,A.J. Willis,2012-12-06 The formative ideas for this symposium originated in 1978 at the IAU Symposium No. 83 on Mass Loss and Evolution of O-type Stars held at Qualicum Beach, Vancouver Island, Canada - WR stars generally figure prominently in O-star meetings and vice versa! Following general approval by the IAU Executive

Committee the initial ideas were cemented at a subsequent meeting, IAU Colloquium No. 59 on The Effects of Mass Loss on Stellar Evolution, held at Miramare, Trieste, Italy in 1980, which was attended by the majority of the present Scientific Organising Committee and at which meeting the outline programme for this symposium was formulated. 1981 was considered an appropriate year in which to hold a meeting on WR stars, since the last IAU Symposium devoted to this stellar class had been held a decade earlier, in Buenos Aires (IAU Symposium No. 49), and during this intervening period a wealth of new observational material had been obtained for WR stars together with significant advances on the theoretical front. The venue for this symposium was chosen from the requirement, which can be inferred from the above, that a meeting on 'hot' stars take place in an appropriate, sunny climate and followed upon the excellent suggestion of Dr. C. Firmani to hold the symposium in Mexico.

Mass Loss and Evolution of O-Type Stars P.S. Conti, C. de Loore, 2012-12-06 The organization of this Symposium had its beginnings at the International Astronomical Union General Assembly in Grenoble in 1976. The initial rounding up of the Scientific Organizing Committee was begun by Drs. Snow and Swings; most of us who became the eventual organizing committee met a few times during the Assembly and formulated the essential outlines of the meeting. Extensive correspondence with all the committee subsequently established the program. The idea was to bring together both observers and theoreticians to discuss the stellar winds and mass loss rates and their effects on evolutions of O-type stars. On the observational side, there are now spectroscopic data from the far UV to the near IR regions concerning the stellar winds. There is also information about the free-free emission in the wind from

the IR and radio portions of the spectrum. Fortunately, these different detection methods give more or less the same mass loss rate for the one star, θ Pup which has been observed at all wavelengths. One of the intents of the first three sessions of this Symposium is to outline the existing data on mass loss rates as it pertains to the O-type stars.

Wolf-Rayet and High-Temperature Stars M.K.V. Bappu, J. Sahade, 2012-12-06 We have in this volume, compiled a connected account of the proceedings of the Symposium on Wolf-Rayet and High-Temperature Stars held at Buenos Aires. The Organizing Committee had assigned broad areas of topical interest to be reviewed by invited speakers. Each of these presentations was followed by lengthy discussions that were tape recorded and transcribed later. These discussions have been edited only to a limited extent. We have shortened them and rearranged them to bring about a greater coherence. We have, however, attempted to retain the tenor of the discussions, the flavour of impromptu remarks and the continuity of an argument. Much of the success of such a venture depends on the contributors to the discussions. To be able to make these thoughts available to a larger audience has been the task of those responsible for the elaborate tape recording of the proceedings. We thank those at the Instituto de Astronomia y Fisica del Espacio for the efficient way in which this responsibility has been discharged. Many at Buenos Aires and Kodaikanal have contributed efficient assistance to the preparation of this volume and we are deeply indebted for their help. In particular, two amongst these, Nora Martinez and A. M. Batcha have contributed overwhelmingly both to the organization of the symposium and the final preparation of the symposium volume. Financial support for this symposium came from the International Astronomical

Union and the Argentine National Research Council.

The Brightest Binaries Dany Vanbeveren, W. van Rensbergen, C. de Loore, 1998-07-31 The evolution of galaxies is governed mainly by the evolution of massive stars whereas the evolution of a massive star depends primarily on its mass, chemical composition and on whether or not the star is a Single object or a binary component. To study the evolution of galaxies, it is therefore essential to know how stellar masses are distributed at birth, how many stars are formed in binaries, and what the mass ratio and orbital period distribution of binaries look like. Massive stars are intrinsically the brightest stars, so that it may be possible to discover their properties in distant groups provided that large telescopes can be used for basic stellar observations. However, until now the observations of massive stars have been reasonably complete only for a small region of our own Galaxy (~ 3 kpc from the Sun). One hopes that the conclusions resulting from these observations hold for the whole Galaxy, for the whole cosmos. With 'The Brightest Stars' of De Jager (1980) in mind, the present monograph is an addendum and an update in which we discuss the observations of 'The Brightest Binaries' in the framework of stellar evolution. A small or intermediate mass star close to the Sun may look brighter than a massive one far away. However, within volume limited star samples, the massive stars are on average also the brightest ones. In the present monograph (similarly as in the work of De Jager), bright means massive. The book consists of four main chapters.

Scientific and Technical Aerospace Reports, 1991

NASA Newsletter for International Ultraviolet Explorer (IUE) United States. National Aeronautics and Space Administration, 1982

Code of Federal Regulations, 2013-03

Luminous Stars and Associations in Galaxies

International Astronomical Union. Symposium, 1986-06-30
Proceedings of the 116th Symposium of the International
Astronomical Union, held at Porto Heli, Greece, May 26-31,
1985

*Executive Documents of the State of Minnesota for the
Year ... Minnesota, 1904*

The Early Type Stars A.B. Underhill, 2012-12-06 At this time when astronomers are being surprised by the discovery of objects which emit a fabulously large amount of energy, that is the quasi-stellar radio sources and the quasi-stellar galaxies, and when by the means of space vehicles X rays, gamma rays and cosmic rays are being observed to come from the depths of interstellar space, one may ask why write a book about stars. Stars seem to be almost incidental when one looks at the universe in terms of exceedingly great energies. Nevertheless, stars exist. They are accessible to study and they have not yet revealed all their secrets. This is enough to arouse interest and to cause one to try to find answers to the questions which arise. The early type stars are particularly interesting because they are spendthrift stars pouring out their energy at a great rate. But their brilliance is also their undoing. They must evolve rather quickly, on an astrophysical scale. Thus by studying these stars we are studying a population in change. The implications from the local and from the cosmological viewpoint are important if one wishes to understand the details of stellar evolution and of galactic structure. Perhaps one of the simplest reasons for writing a book about the early type stars is to see if some of the conundrums presented by the spectra of these stars can be unravelled when all the available information is brought together.

Title 30 Mineral Resources Parts 200 to 699 (Revised

as of July 1, 2013) Office of The Federal Register, Enhanced by IntraWEB, LLC, 2014-07-01 The Code of Federal Regulations Title 30 contains the codified United States Federal laws and regulations that are in effect as of the date of the publication pertaining to U.S. mineral resources, including: coal mining and mine safety; surface mining, fracking and reclamation; offshore oil, gas and sulphur drilling, safety, oil spills response; minerals leasing and revenues from public lands.

Proceedings of the ... Liège International Astrophysical Colloquium ,1996

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