Disk

G. Belvedere

Server Disk Management in a Windows Environment Drew Robb,2003-09-26 Hard drives and disk management receive scant attention from the industry press, yet recent surveys have identified disk failure as the #1 source of server downtime. Combine this fact with the skyrocketing TCO of data storage management, and it is apparent that server disk management is a subject deserving of much more scrutiny. Server Disk

Dvorak's Inside Track to the Mac John Dvorak,1992 John Dvorak and his co-authors have compiled the definitive, information-packed book/disk package on the Macintosh, loaded with all the insights and tricks that Mac users could ever want. From Mac components to operating systems, multimedia, graphics, and desktop publishing, you'll find it here. (Apple/Macintosh)

Comparison of Methods to Determine Disk and Heartwood Areas Michael Carl Wiemann,2002

IBM DS8880 Product Guide (Release 8.51) Bert Dufrasne,Peter Kimmel,Stephen Manthorpe,Tamas Toser,IBM Redbooks,2019-01-02 This IBM Redbooks® Product Guide gives an overview of the features and functions that are available with the IBM DS8880 models running microcode Release 8.51 (DS8000 License Machine Code 8.8.51.xx.xx). The IBM DS8880 architecture relies on powerful IBM POWER8® processor-based servers that manage the cache to streamline disk input/output (I/O), maximizing performance and throughput. These capabilities are further enhanced with the availability of the second generation of high-performance flash enclosures (HPFE Gen-2). The IBM DS8888, DS8886, and DS8884 models excel at supporting the IBM Z Enterprise server and IBM Power server environments, offering many synergy features.

Hard Disk Drive Servo Systems Ben M. Chen,Tong Heng Lee,Kemao Peng,Venkatakrishnan Venkataramanan,2006-06-09 The series Advances in Industrial Control aims to report and encourage technology transfer in control engineering. The rapid development of control technology has an impact on all areas of the control discipline. New theory, new controllers, actuators, sensors, new industrial processes, computer methods, new applications, new philosophies , new challenges. Much of this development work resides in industrial reports, feasibility study papers and the reports of advanced collaborative projects. The series offers an opportunity for researchers to present an extended exposition of such new work in all aspects of industrial control for wider and rapid dissemination. Hard disk drive systems are ubiquitous in today’s computer systems and the technology is still evolving. There is a review of hard disk drive technology and construction in the early pages of this monograph that looks at the characteristics of the disks and there it can be read that: “bit density... continues to increase at an amazing rate”, “spindle speed... the move to faster and faster spindle speeds continue”, “form factors... the trend...is downward... to smaller and smaller drives”, “performance... factors are improving”, “redundant arrays of inexpensive disks... becoming increasingly common, and is now seen in consumer desktop machines”, “reliability... is improving slowly... it is very hard to improve the reliability of a product when it is changing rapidly” and finally “interfaces... continue to create new and improved standards... to match the increase in performance of the hard disks themselves”.

Hard Disk Management Thomas Cain,Nancy Woodard Cain,1990 Gives the essentials of DOS ; a complete overview of the DOS 4.0 shell utility program

; detailed information on disks, disk drives, and disk-related commands ; programming with DOS ; advanced features of DOS.

Redundant Disk Arrays Garth A. Gibson,1992 Disk arrays, coupled with emerging small disk technology, promise to provide a badly needed increase in the performance of secondary storage systems. Because high failure rates arise with a large number of disks, however, simple redundancy schemes are used to ensure data reliability. This monograph investigates the data encoding, performance, and reliability of redundant disk arrays. Gibson reviews the performance advantages of striping data across multiple disks, evaluates the performance lost to the maintenance of redundant data, provides evidence that disk lifetimes can be modeled as exponential random variables, and develops and applies analytic models of data reliability in redundant disk arrays suffering dependent failure modes and featuring on-line spare disks. Garth A. Gibson is a Research Computer Scientist in the School of Computer Science at Carnegie-Mellon University.

Theory of Accretion Disks F. Meyer,Wolfgang J. Duschl,Juhan Frank,Emmi Meyer-Hofmeister,2012-12-06 With the advent of space observatories and modern developments in ground based astronomy and concurrent progress in the theoretical understanding of these observations it has become clear that accretion of material on to compact objects is an ubiquitous mechanism powering very diverse astrophysical sources ranging in size and luminosity by many orders of magnitude. A problem common to these systems is that the material accreted must in general get rid of its angular momentum and this leads to the formation of an Accretion Disk which allows angular momentum re-distribution and converts potential energy into radiation with an efficiency which can be higher than the nuclear burning yield. These systems range in size from quasars and active galactic nuclei to accretion disks around forming stars and the early solar system and to compact binaries such as cataclysmic variables and low-mass X-ray binaries. Other objects that should be mentioned in this context are 88433, the black hole binary candidates, and possibly gamma-ray burst sources. Observations of these systems have provided important constraints for theoretical accretion disk models on widely differing scales, lumi nosities, mass-transfer rates and physical environments.

Circumstellar Dust Disks and Planet Formation Roger Ferlet,Alfred Vidal-Madjar,1994

Oracle Disk I/O Tuning Mike Ault,2004-04 Covering all aspects of Oracle disk I/O tuning, this book explores disk performance, RAID management, Oracle data file performance, and Oracle data segment internals. Also explored is physical disk I/O, which includes disk device internals, detecting disk bottlenecks, disk organization techniques, and disk striping and disk load balancing. Highlighted are RAID and Oracle performance as well as techniques for effective use of RAID with Oracle. Additionally, Oracle data file internals are considered and how to use multiple data block sizes to detect and repair Oracle data segment bottlenecks and segment waits is described. Oracle segment management is illustrated, and the effective use of Oracle segment partitioning, segment slot internals, and monitoring segment I/O is explained.

Physical Processes in Circumstellar Disks Around Young Stars Paulo J. V. Garcia,2011-05-15 Circumstellar disks are vast expanses of dust that form around new stars in the earliest stages of their birth. Predicted by

astronomers as early as the eighteenth century, they weren't observed until the late twentieth century, when interstellar imaging technology enabled us to see nascent stars hundreds of light years away. Since then, circumstellar disks have become an area of intense study among astrophysicists, largely because they are thought to be the forerunners of planetary systems like our own—the possible birthplaces of planets. This volume brings together a team of leading experts to distill the most up-to-date knowledge of circumstellar disks into a clear introductory volume. Understanding circumstellar disks requires a broad range of scientific knowledge, including chemical processes, the properties of dust and gases, hydrodynamics and magnetohydrodynamics, radiation transfer, and stellar evolution—all of which are covered in this comprehensive work, which will be indispensable for graduate students, seasoned researchers, or even advanced undergrads setting out on the study of planetary evolution.

From Protoplanetary Disks to Planet Formation Philip J. Armitage,Wilhelm Kley,2019-02-02 Is the Sun and its planetary system special? How did the Solar system form? Are there similar systems in the Galaxy? How common are habitable planets? What processes take place in the early life of stars and in their surrounding circumstellar disks that could impact whether life emerges or not? This book is based on the lectures by Philip Armitage and Wilhelm Kley presented at 45th Saas-Fee Advanced Course „From Protoplanetary Disks to Planet Formation" of the Swiss Society for Astrophysics and Astronomy. The first part deals with the physical processes occurring in proto-planetary disks starting with the observational context, structure and evolution of the proto-planetary disk, turbulence and accretion, particle evolution and structure formation. The second part covers planet formation and disk-planet interactions. This includes in detail dust and planetesimal formation, growth to protoplanets, terrestrial planet formation, giant planet formation, migration of planets, multi-planet systems and circumbinary planets. As Saas-Fee advanced course this book offers PhD students an in-depth treatment of the topic enabling them to enter on a research project in the field.

Dust-Gas Instabilities in Protoplanetary Disks Ryosuke Tominaga,2022-04-25 How planets form is one of the long-standing questions in astrophysics. In particular, formation scenarios of planetesimals which are kilometer-sized bodies and a precursor of planets are still unclear and under debate although some promising mechanisms have been proposed. This book highlight disk instabilities that have the potential to explain the origin of planetesimals. Using linear analyses and numerical simulations, it addresses how a disk evolves through the development of instabilities, and also presents a new instability driven by dust coagulation. As a result, the simulation demonstrates a scenario of planetesimal formation: A successive development of multiple instabilities triggers planetesimal formation in resulting dusty rings.

High Angular Resolution Studies of the Structure and Evolution of Protoplanetary Disks Joshua Eisner,2005 Young stars are surrounded by massive, rotating disks of dust and gas, which supply a reservoir of material that may be incorporated into planets or accreted onto the central star. In this dissertation, I use high angular resolution observations at a range of wavelengths to understand the structure, ubiquity, and evolutionary timescales of protoplanetary disks. First, I describe a study of Class I

protostars, objects believed to be at an evolutionary stage between collapsing spherical clouds and fully-assembled young stars surrounded by protoplanetary disks. I use a Monte Carlo radiative transfer code to model new 0.9 micron scattered light images, 1.3 mm continuum images, and broadband spectral energy distributions. This modeling shows that Class I sources are probably surrounded by massive protoplanetary disks embedded in massive infalling envelopes. For the best-fitting models of the circumstellar dust distributions, I determine several important properties, including envelope and disk masses, mass infall rates, and system inclinations, and I use these results to constrain the evolutionary stage of these objects. Second, I discuss observations of the innermost regions of more evolved disks around T Tauri and Herbig Ae/Be stars, obtained with the Palomar Testbed and Keck Interferometers. I constrain the spatial and temperature structure of the circumstellar material at sub-AU radii, and demonstrate that lower-mass stars are surrounded by inclined disks with puffed-up inner edges 0.1-1 AU from the star. In contrast, the truncated inner disks around more massive stars may not puff-up, indicating that disk structure depends on stellar properties. I discuss the implications of these results for disk accretion, terrestrial planet formation and giant planet migration. Finally, I put these detailed studies of disk structure into a broader context by constraining the mass distribution and evolutionary timescales of circumstellar disks. Using the Owens Valley Millimeter Array, I mapped the millimeter continuum emission toward >300 low-mass stars in the NGC 2024 and Orion Nebula clusters. These observations demonstrate that the average disk mass in each cluster is comparable to the minimum-mass protosolar nebula, and that there may be disk evolution on one million year timescales.

Server Disk Management in a Windows Environment Drew Robb,2003-09-26 Hard drives and disk management receive scant attention from the industry press, yet recent surveys have identified disk failure as the #1 source of server downtime. Combine this fact with the skyrocketing TCO of data storage management, and it is apparent that server disk management is a subject deserving of much more scrutiny. Server Disk Management in a Windows Environment explains the basic elements of disks and disk architectures, and explores how to successfully manage and maintain functionality within a Windows environment. The author focuses on critical issues that are often ignored by other books on this subject, issues including disk quotas, fragmentation, optimization, hard drive reliability, asset management, software deployment, and system forensics. This book is a valuable resource for IT veterans and those who are less tech-savvy. Students and IT newcomers can access multiple definitions and examples of fundamentals such as disk organization, file systems, and basic maintenance actions. Chapters examining hardware and software management and industry trends provide further insight into how you can maintain and monitor disks, and have them perform at maximum efficiency.

Theory of Accretion Disks 2 Wolfgang J. Duschl,Juhan Frank,F. Meyer,Emmi Meyer-Hofmeister,Werner M. Tscharnuter,2012-12-06 Accretion disks in astrophysics represent the characteristic flow by which compact bodies accrete mass from their environment. Their intrinsically high luminosity, and recent progress in observational accessibility at all wavelength bands, have led to rapidly growing awareness of their importance and made them the object of intense research on widely different scales, ranging from binary stars to

young stellar objects and active galactic nuclei. This book contains the proceedings of the NATO Advanced Workshop on `Theory of Accretion Disks 2' for which some of the most active researchers in the different fields came together at the Max-Planck-Institut for Astrophysics in Garching in March, 1993. Its reviews and contributions give an up-to-date account of the present status of our understanding and provide a stimulating challenge in discussions of open questions in a rapidly developing field.

Warped Disks and Inclined Rings Around Galaxies Stefano Casertano,Penny D. Sackett,Franklin H. Briggs,1991-05-30 Most galaxies are in clusters, where tidal interactions are not uncommon. Tidal and dynamical interaction in galaxies are of importance in studying evolution. A large amount of data has been collected on dust-lane ellipticals, polar ring galaxies, spirals with extended warps, and galaxies with inclined HI rings or unusual 'tails'. This book is a record of a meeting which was held at the University of Pittsburgh. It provided an informal, yet focused environment for the interaction of astronomers who have addressed these questions with a wide variety of skills, techniques and points of view.

IBM i 6.1 Independent ASPs: A Guide to Quick Implementation of Independent ASPs Scott Vetter,James Baer,Julie Cantrell,Harlon Trowbridge,IBM Redbooks,2009-12-10 This IBM® Redbooks® publication explains how to configure and manage independent disk pool (IASP) functionality of IBM i 6.1. It is designed to help IBM technical professionals, business partners, and customers understand and implement independent disk pools in the IBM i 6.1. In addition, this publication provides the background information that is necessary to plan, implement, and customize this functionality to your particular environment. It provides guidance on running user applications with either application data or most application objects residing in an independent disk pool. Considering that you can also use independent disk pools in a cluster environment, this publication shows you the basic steps to make your independent disk pool switchable between two Power SystemsTM servers or a single server with multiple LPARs. Independent auxiliary storage pools have many business and technical advantages for Power Systems using IBM i. Not only are independent auxiliary storage pools (IASPs) easy to create and maintain, most applications can use them by simple work management changes. IASPs can provide immediate benefits to your enterprise.

Accretion Disks and Magnetic Fields in Astrophysics G. Belvedere,2012-12-06 Proceeding of the European Physical Society Study Conference, held in Noto (Sicily), Italy, June 16-20, 1988

Structure and Emission Properties of Accretion Disks C. BERTOUT (Ed),1991

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